## AERONAUTICAL RESEARCH COUNCIL REPORTS AND MEMORANDA

# Forced Flow against a <br> Rotating Disc 

By
Miss D. M. Hannah, B.A.

Crover Copyright Reserved

LONDON: HER MAJESTY'S STATIONERY OFFICE '1952

FIVE SHILLINGS NET

# Forced Flow against a Rotating Disc 

By<br>Miss D. M. Hannah, B.A.

Communicated by Prof. W. G. Bickley

Reports and Memoranda No. 2772
April, 1947

1. Summary.-The steady motion of an incompressible viscous fluid due to an infinite rotating plane lamina has been considered by Von Kármán ${ }^{1}$ and by Cochran ${ }^{2}$ : the motion of fluid flowing with axial symmetry towards an infinite stationary plane lamina has been dealt with by Hömann ${ }^{3}$. The present paper deals with the general question of steady irrotational flow with axial symmetry against an infinite rotating lamina, of which the above are two special cases.
2. Notation.-Let $\gamma, \theta, z$ be cylindrical polar co-ordinates: the lamina is taken to be the plane $z=0$, and is rotating with constant angular velocity $\omega$ about the axis $\gamma=0$. Let $u, v$, $w$ be the components of fluid velocity in the directions of $\gamma, \theta, z$, increasing, and let $p$ be the pressure. Axial symmetry is assumed, so that all quantities are independent of $\theta$.
3. Solution for Non-Viscous Fluid.-When viscosity is ignored, the solution will be the same - for a rotating as for a stationary disc, since the rotation of the disc only affects the fluid velocity through the viscous drag which it exerts. This solution may then be obtained from that for a source in the presence of an infinite plane by letting the distance between source and plane tend to infinity.

If the source is of strength $m$, at the point $(0,0, a)$, and the plane is given by $z=0$, then it is known that the potential of the resulting flow is

$$
\begin{equation*}
\phi=m\left(\frac{1}{l}+\frac{1}{l^{\prime}}\right) \quad . . \quad . . \quad . . \quad . . \quad \therefore \quad \quad . \quad . . \quad . \tag{3.1}
\end{equation*}
$$

where

$$
\left.\begin{array}{rl}
l & =\left\{(a-z)^{2}+r^{2}\right\}^{1 / 2}  \tag{3.2}\\
l^{\prime} & =\left\{(\dot{a}+z)^{2}+r^{2}\right\}^{1 / 2}
\end{array}\right\} \quad \begin{array}{llll}
\ldots & \ldots & \ldots & \ldots
\end{array}
$$

Expanding the expression for $\phi$ in powers of $1 / a$, we obtain

$$
\begin{equation*}
\phi=\frac{m}{a}\left\{2+\frac{2 z^{2}-r^{2}}{2 a^{2}}+\mathrm{O}\left(\frac{1}{a^{3}}\right)\right\} . \quad . \quad . . \quad . \quad . . \tag{3.3}
\end{equation*}
$$

So that if $m$ and $a$ tend to infinity in such a way that $m / a^{3}$ remains equal to $k$, a finite constant, then

$$
\begin{equation*}
\phi=k\left(z^{2}-\frac{1}{2} r^{2}\right) \quad \text {.. } \quad . \quad \text {.. .. .. .. .. .. } \tag{3.4}
\end{equation*}
$$

ignoring the constant term $2 k a^{2}$.

From this,

$$
\left.\begin{array}{l}
u=-\frac{\partial \phi}{\partial r}=k r  \tag{3.5}\\
w=-\frac{\partial \phi}{\partial z}=-2 k z
\end{array}\right\} \quad \ldots \quad . . \quad . . \quad \ldots \quad \ldots
$$

and by Bernouilli's equation

$$
\int \frac{d p}{\rho}=-\frac{1}{2} q^{2}=-\frac{1}{2} k^{2}\left(\gamma^{2}+4 z^{2}\right)
$$

so that

$$
\begin{equation*}
\frac{1}{\rho}\left(p_{0}-p\right)=\frac{1}{2} k^{2}\left(r^{2}+4 z^{2}\right) \quad . \quad . \quad . \quad . \quad . \quad . \quad . \quad . \tag{3.6}
\end{equation*}
$$

where $p_{0}$ is the pressure at the stagnation point, i.e., at the origin.
This is the complete solution for the non-viscous case ; the result is quoted by Hömann. We shall use it as giving a guide to the form of the solution when viscosity is taken into account.
4. Equations of Motion for Viscous Fluid.-The general equations governing the steady motion of the fluid are, assuming axial symmetry
continuity

$$
\begin{equation*}
u+r \frac{\partial u}{\partial r}+r \frac{\partial w}{\partial z}=0 \quad \text {.. .. .. .. .. .. .. } \tag{4.1}
\end{equation*}
$$

Navier-Stokes

$$
\begin{align*}
& u \frac{\partial u}{\partial r}+w \frac{\partial u}{\partial z}-\frac{v^{2}}{r}=-\frac{1}{\rho} \frac{\partial p}{\partial r}+\nu\left(\frac{\partial^{2} u}{\partial r^{2}}+\frac{1}{r} \frac{\partial u}{\partial r}+\frac{\partial^{2} u}{\partial z^{2}}-\frac{u}{r^{2}}\right) \ldots  \tag{4.2}\\
& u \frac{\partial v}{\partial r}+w \frac{\partial v}{\partial z}+\frac{u v}{r}=v\left(\frac{\partial^{2} v}{\partial r^{2}}+\frac{1}{r} \frac{\partial v}{\partial r}+\frac{\partial^{2} v}{\partial z^{2}}-\frac{v}{r^{2}}\right)  \tag{4.3}\\
& \ldots \ldots  \tag{4.4}\\
& u^{2} \frac{\partial w}{\partial r}+w \frac{\partial w}{\partial z}=-\frac{1}{\rho} \frac{\partial p}{\partial z}+v\left(\frac{\partial^{2} w}{\partial r^{2}}+\frac{1}{r} \frac{\partial w}{\partial r}+\frac{\partial^{2} w}{\partial z^{2}}\right) \\
& \ldots \\
& \hline
\end{align*}
$$

For boundary conditions, we have that at the disc the velocities of fluid and disc are equal, i.e.,

$$
\begin{equation*}
u=0, v=\omega r, w=0 \text { when } z=0 \tag{4.5}
\end{equation*}
$$

And far from the disc the velocity will tend to that attained in the non-viscous case, at least for $u$ and $v$ : there is a finite change in $w$, the axial inflow velocity, due to the radial outflow now caused by the disc's rotation, i.e.,

$$
\begin{equation*}
u \rightarrow k r, v \rightarrow 0, w \rightarrow-2 k z-c \text { as } z \rightarrow \infty \tag{4.6}
\end{equation*}
$$

where $c$ is some constant to be determined.
5. We seek a solution in the form

$$
u=\gamma u_{1}(z), v=\gamma v_{1}(z), z v=w_{1}(z),
$$

$$
\begin{equation*}
\frac{1}{\rho}\left(p_{0}-p\right)=\frac{1}{2} k^{2} r^{2}+p_{1}(z) \quad . \quad . \quad . \quad . \quad \text {.. .. } \tag{5.1}
\end{equation*}
$$

$$
\begin{align*}
2 u_{1}+\frac{d w_{1}}{d z} & =0 \tag{5.2}
\end{align*} \cdots \quad \ldots .
$$

with boundary conditions

$$
\left.\begin{array}{c}
u_{1}=0, \quad v_{1}=\omega, \quad w_{1}=0 \text { when } z=0  \tag{5.6}\\
u_{1} \rightarrow k, \quad v_{1} \rightarrow 0, \quad w_{1} \rightarrow-2 k z-c \text { as } z \rightarrow \infty
\end{array}\right\} \quad \ldots \quad \ldots \quad .
$$

6. The equations may now be made non-dimensional by the substitutions

$$
\begin{array}{ccc}
u_{1}=\eta N^{\prime}(\zeta), v_{1}=\omega G(\zeta), w_{1}=-2(\nu \eta)^{1 / 2} & N(\zeta) \\
p_{1}=\nu \eta P(\zeta) & \cdots & \cdots \tag{6.1}
\end{array} .
$$

where

$$
\eta=\left(k^{2}+\omega^{2}\right)^{1 / 2}, \zeta=(\eta / v)^{1 / 2} z .
$$

Equation (5.2) is then automatically satisfied, and (5.3) (5.4) (5.5), become

$$
\begin{array}{lccclll}
N^{\prime 2}-\Omega^{2} G^{2}-2 N N^{\prime \prime}=K^{2}+N^{\prime \prime \prime} & . & . & . & . \\
2 G N^{\prime}-2 G^{\prime} N=G^{\prime \prime} & . & . & . & . & . & . \\
4 N N^{\prime}-2 N^{\prime \prime}=P^{\prime} & . & . & . & . & . & .
\end{array}
$$

where $\Omega=\omega / \eta, K=k / \eta$ so that $\Omega^{2}+K^{2}=1$ and primes denote differentiation with respect to $\zeta$. Boundary conditions now are

$$
\left.\begin{array}{l}
N=0, \quad N^{\prime}=0, \quad G=1 \text { when } \zeta=0 .  \tag{6.5}\\
N \rightarrow K \zeta+\frac{1}{2} C, \quad N^{\prime} \rightarrow K, \quad G \rightarrow 0 \text { as } \zeta \rightarrow \infty
\end{array}\right\} \quad . \quad .
$$

where $C=c / \sqrt{ }(\nu / \eta)$
These substitutions, and the resulting equations, reduce with slight modifications to those used by von Kármán and Cochran when $k=0$, and to those used by Hömann when $\omega=0$. There is however an error of sign in Cochran's equation (10) which corresponds to (6.4) above ; the term corresponding to $P^{\prime}$ is given a negative sign. This affects the integral of the equation also.
7. Solution of the Equations.-Equation (6.4) integrates immediately to give

$$
\begin{equation*}
P=2\left(N^{2}+N^{\prime}\right) \tag{7.1}
\end{equation*}
$$

since $P=0$ when $N, N^{\prime}$ vanish.
Equations (6.2), (6.3) have been integrated numerically for various given values of $\mu=\omega / k$. The method was to find a series solution for $N, G$ which held for small values of $\zeta$ : this involves two unknown parameters $a, b$. By successive use of forward integration formulæ on values calculated from these series using trial values of $a, b$, tables of values of the functions were built up for increasing $\zeta$. Then by adjusting the trial values, solutions could be found which satisfied the boundary conditions for large $\zeta$. Finally these solutions were keyed in with values computed from an asymptotic expansion for large $\zeta$, and the three parameters of this expansion were thereby determined.
8. For small $\zeta$ the series expansions are

$$
\begin{align*}
& N=a \zeta^{2}-\frac{1}{6} \zeta^{3}-\frac{1}{12}\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b \zeta^{4}-\frac{1}{60}\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b^{2} \zeta^{5} \\
& N^{\prime}=2 a \zeta-\frac{1}{2} \zeta^{2}-\frac{1}{3}\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b \zeta^{3}-\frac{1}{12}\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b^{2} \zeta^{4} \\
& +\frac{1}{30}\left(\frac{1-\mu^{2}}{1+\mu^{2}}\right) a \zeta^{5} \\
& N^{\prime \prime}=2 a-\zeta-\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b \zeta^{2}-\frac{1}{3}\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b^{2} \zeta^{3}+\frac{1}{6}\left(\frac{1-\mu^{2}}{1+\mu^{2}}\right) a \zeta^{4} \\
& +\frac{1}{60}\left(\frac{\mu^{2}-1-8 \mu^{2} a b}{1+\mu^{2}}\right) \zeta^{5} \\
& G=1+b \zeta+\frac{2}{3} a \zeta^{3}+\frac{1}{12}(2 a b-1) \zeta^{4}-\frac{1}{30}\left(\frac{1+2 \mu^{2}}{1+\mu^{2}}\right) \zeta^{5} \\
& G^{\prime}=b+2 a \zeta^{2}+\frac{1}{3}(2 a b-1) \zeta^{3}-\frac{1}{6}\left(\frac{1+2 \mu^{2}}{1+\mu^{2}}\right) \zeta^{4} \\
& -\frac{2}{15}\left\{2 a^{2}+\left(\frac{\mu^{2}}{1+\mu^{2}}\right) b^{2}\right\} \zeta^{5} \quad \cdots \quad . . \quad . \quad . \quad . \tag{8.1}
\end{align*}
$$

The asymptotic solution for large $\zeta$ is found by putting

$$
\left.\begin{array}{ll}
N=K \zeta+\frac{1}{2} C+x & \text { i.e., } N^{\prime}=K+x^{\prime}, N^{\prime \prime}=x^{\prime \prime}, N^{\prime \prime \prime}=x^{\prime \prime \prime}  \tag{8.2}\\
G=y & \text { i.e., } G^{\prime}=y^{\prime}, G^{\prime \prime}=y^{\prime \prime}
\end{array}\right\}
$$

where $x$ and $y$ are small and powers higher than the first are neglected. Substituting these in (6.2), (6.3)

$$
\begin{aligned}
& x^{\prime \prime \prime}+2\left(K \zeta+\frac{1}{2} C\right) x^{\prime \prime}-2 K x^{\prime}=0 \\
& y^{\prime \prime}+2\left(K \zeta+\frac{1}{2} C\right) y^{\prime}-2 K y=0
\end{aligned}
$$

i.e., $x^{\prime}, y$ both satisfy

$$
\begin{equation*}
f^{\prime \prime}+2\left(K \zeta+\frac{1}{2} C\right) f^{\prime}-2 K f=0 . \quad . . \quad . . \quad . . \quad . . \quad . \tag{8.3}
\end{equation*}
$$

If now $K \neq 0$, write $K \zeta+\frac{1}{2} C=t$ and (8.3) becomes

$$
\begin{equation*}
K \frac{d^{2} f}{d t^{2}}+2 t \frac{d f}{d t}-2 f=0 . \quad . . \quad . \quad . . \quad . . \quad . . \tag{8.4}
\end{equation*}
$$

The complete solution of this is

$$
\begin{equation*}
f=R_{1} t+R_{2} t \int^{t} u^{-2} \mathrm{e}^{-w^{2} / \mathbb{}} d u \quad . . \quad . . \quad . \quad \quad \cdots \quad . . \quad . \tag{8.5}
\end{equation*}
$$

where $R_{1}, R_{2}$ are arbitrary constants. To satisfy the condition that $f$ tends to zero as $t$ tends to infinity, $R_{1}$ must vanish, and so the solutions for $x^{\prime}$ and $y$ are

$$
\left.\left.\begin{array}{l}
x^{\prime}=A t \int_{i}^{\infty} u^{-2} \mathrm{e}^{-u^{2} / K} d u  \tag{8.6}\\
y=B t \int_{t}^{\infty} u^{-2} \mathrm{e}^{-u^{2} / K} d u
\end{array}\right\} \quad \ldots \quad \begin{array}{llll}
\end{array}\right\} \quad \ldots \quad . .
$$

where $A, B$ are arbitrary constants. Since

$$
\left.N=K \zeta+\frac{1}{2} C+\text { (terms of } 1 \text { st order }\right)
$$

we may replace $K \zeta+\frac{1}{2} C(=t)$ by $N$ in the small terms, and finally, simplifying the integral, write the asymptotic solution for $N^{\prime}, G$ as

$$
\left.\begin{array}{l}
N^{\prime} \sim K-A\left\{\mathrm{e}^{-N^{2} / K}-2 N / \sqrt{ } K \int_{N / V K}^{\infty} \mathrm{e}^{-u^{2}} d u\right\} \\
G \sim B\left\{\mathrm{e}^{-N^{2} / K}-2 N / \sqrt{ } K \int_{N / / K}^{\infty} \mathrm{e}^{-u^{2}} d u\right\} \tag{8.7}
\end{array}\right\}
$$

If now $K=0,(8.3)$ becomes

$$
f^{\prime \prime}+C f^{\prime}=0
$$

so that $x^{\prime}, y$ are both proportional to $\mathrm{e}^{-c \xi_{\xi}}$, and the asymptotic solutions are

$$
\left.\begin{array}{rl}
N^{\prime} & \sim A_{0} \mathrm{e}^{-C \xi}  \tag{8.8}\\
G & \sim B_{0} \mathrm{e}^{-C 5}
\end{array}\right\} \quad \ldots \quad . . \quad . \quad . \quad . \quad .
$$

9. It is of interest to consider the relationship between the two solutions (8.7), (8.8). Both satisfy equation (8.3)

$$
f^{\prime \prime}+2\left(K \zeta+\frac{1}{2} C\right) f^{\prime}-2 K f=0
$$

with formal boundary conditions

$$
\begin{aligned}
& f(K, \zeta)=\phi_{0}(K) \text { when } \zeta=\zeta_{0} \\
& f(K, \zeta) \rightarrow 0 \quad \text { when } \zeta \rightarrow \infty
\end{aligned}
$$

where $\phi_{0}(K)$ is the value obtained from the integrated solution with which a join is to be made at $=\zeta_{0}$. Solutions of (8.3) are

$$
\begin{aligned}
& f(K, \zeta)=K \zeta+\frac{1}{2} C \\
& f(K, \zeta)=\left(K \zeta+\frac{1}{2} C\right) \int_{K \zeta++C}^{\infty} u^{-2} \mathrm{e}^{-u^{2} / K} d u
\end{aligned}
$$

so that the characteristic equation is
$\underset{\zeta}{\mathrm{L} t} \rightarrow \infty\left|\begin{array}{cc}K \zeta+\frac{1}{2} C & K \zeta_{0}+\frac{1}{2} C \\ 0 & \left(K \zeta_{0}+\frac{1}{2} C\right) \int_{K \zeta_{0}+\frac{1}{4} C}^{\infty} u^{-2} \mathrm{e}^{-u^{2} / K} d u\end{array}\right|=0 . \quad \ldots \quad .$.
This has no root for $0 \leqslant K \leqslant 1$, i.e., there is no characteristic number for $K$ in this range. Then, by a known existence theorem, since the coefficients of equation (8.3) and the boundary conditions are all uniformly continuous functions of $\zeta$ and analytic in $K$ the solution is also uniformly continuous in and analytic in $K$ for $0 \leqslant K \leqslant 1$. In particular

$$
f(K, \zeta) \rightarrow f(0, \zeta) \text { as } K \rightarrow 0
$$

so that

$$
\begin{equation*}
A\left\{\mathrm{e}^{-N^{2} / K}-\frac{2 N}{\sqrt{ } K} \int_{N / V K}^{\infty} \mathrm{e}^{-u^{2}} d u\right\} \rightarrow A_{0} \mathrm{e}^{-c \zeta} \text { as } K \rightarrow 0 . \quad . \quad . . \quad . \quad \ldots \tag{9.2}
\end{equation*}
$$

For small $K$

$$
\frac{2 N}{\sqrt{ } K} \int_{N / / K}^{\infty} \mathrm{e}^{-u^{2}} d u \sim \mathrm{e}^{-N^{2} / K}\left(1-\frac{K}{2 N^{2}}+\ldots\right)
$$

Substituting in (9.2)

$$
A \mathrm{e}^{-N^{2} / K} \frac{K}{2 N^{2}} \rightarrow A_{0} \mathrm{e}^{-c \xi}
$$

which gives finally

$$
\begin{equation*}
A \rightarrow A_{0} \frac{C^{2}}{2 K} \mathrm{e}^{\mathrm{C}^{2 / 4 K}}(1+O(K)) . \quad . \quad . \quad . \quad . . \quad . . . \tag{9.3}
\end{equation*}
$$

This equation shows the variation of $A$ with $K$ when $K$ is small. A similar expression holds for $B$.
10. Details of the Numerical Solution.-The solution was obtained for five values of $\mu$ : $\mu=0, \frac{1}{2}, 1,2, \infty$. These correspond respectively to $K=1, \Omega=0 ; K=0 \cdot 894, \Omega=0.447$; $K=0.707, \Omega=0.707 ; K=0.447, \Omega=0.894 ; K=0, \Omega=1$. The case $\mu=0$ had already been treated by Hömann, but his numerical results were inaccurate and have been corrected. Accurate results for the case $\mu=\infty$ had been produced by Cochran, and are quoted here for completeness.

For a given value of $\mu$, trial values $\alpha, \beta$ of $a, b$ could be guessed from the evidence of previous solutions or by rough approximations. Using $\alpha, \beta$, the first few entries in the tables of $N, N^{\prime}, N^{\prime \prime}, N^{\prime \prime \prime}, G, G^{\prime}, G^{\prime \prime}$ were computed from the series expansions (8.1) and the equations (6.2), (6.3). The $\zeta$ interval was taken as $0 \cdot 1$. Forward numerical integration was then performed using the three-strip formula of Milne

$$
y_{4}-y_{0}=\frac{4 h}{3}\left(2 q_{1}-q_{2}+2 q_{3}\right)
$$

checked by Simpson's rule

$$
y_{4}-y_{0}=\frac{2 \hbar}{3}\left(q_{0}+4 q_{2}+q_{4}\right)
$$

where $q=y^{\prime}$ and $h$ is the interval. The errors of these two formule are almost equal and in opposite senses $\left(+14 / 45 q^{\text {iv }} h^{5},-16 / 45 q^{\text {iv }} h^{5}\right)$ so that where the values obtained from them differ, the mean, or more accurately the weighted mean (in the ratio $8: 7$ ) is a more accurate estimate of $y_{k}-y_{0}$ than either separately.

A running check on alternate values using the Simpson two-strip rule in the form

$$
\begin{aligned}
& y_{2 n}-y_{0}=\frac{h}{3}\left(q_{0}+4 q_{1}+2 q_{2}+4 q_{3}+\ldots \ldots+4 q_{2 n-1}+q_{2 n}\right. \\
& y_{2 n+1}-y_{1}=\frac{h}{3}\left(q_{1}+4 q_{2}+2 q_{3}+4 q_{4}+\ldots \ldots+4 q_{2 n}+q_{2 n+1}\right)
\end{aligned}
$$

was carried out on every entry in the $N^{\prime \prime}$ and $G^{\prime}$ columns for $\zeta \geqslant 1.5$; by linking the values obtained with the two entries $y_{0}, y_{1}$ at the very beginning of the table, this check smoothes out most of the tendency to periodic fluctuation which may be introduced by the ordinary method. This fluctuation first occurs in $N^{\prime \prime}$ and in $G^{\prime}$, and if controlled there does not affect the columns produced by integrating these.

The integration was carried on until the functions $N^{\prime}, G$ showed signs of becoming constant, generally at about $\zeta=2 \cdot 0$. By extrapolation an estimate of their limiting values $n_{1}, g_{1}$, could be obtained. Then similar integrations were carried out using trial values ( $\alpha+\varepsilon_{1}, \beta$ ), $\left(\alpha, \beta+\varepsilon_{1}\right),\left(\alpha+\varepsilon_{1}, \beta+\varepsilon_{1}\right)$ where $\left|\varepsilon_{1}\right|$ was taken generally as 0.1 (in one case 0.01 ) and the sign of $\varepsilon_{1}$ was determined by inspection so as to decrease the errors of $n_{1}, g_{1}$. Limits ( $n_{2}, g_{2}$ ), $\left(n_{3}, g_{3}\right),\left(n_{4}, g_{4}\right)$ were obtained for these three sets of trial values, and bi-linear interpolation for the given boundary conditions then led to a new trial starting point $\alpha_{11}, \beta_{11}$. This gave rise to another set of trial values $\left(\alpha_{11}+\varepsilon_{11}, \beta_{11}\right)\left(\alpha_{11}, \beta_{11}+\varepsilon_{11}\right),\left(\alpha_{11}+\varepsilon_{11}, \beta_{11}+\varepsilon_{11}\right)$ where $\left|\varepsilon_{11}\right|$ was $0 \cdot 1\left|\varepsilon_{1}\right|$. From the corresponding set of limits better trial values $\alpha_{111}, \beta_{111}$ could be determined, and in general these were now sufficiently near the true for the required adjustment to find $a, b$ to be obvious.

The process of finding the integrated solutions corresponding to various trial values of the initial parameters could in some cases be shortened. Where numerical solutions $N_{1}, N_{1}{ }^{\prime}$ $G_{1}{ }^{\prime \prime} ; N_{2}, N_{2}{ }^{\prime} \ldots \ldots . G_{2}{ }^{\prime \prime}$ had been worked out for initial values $\alpha_{1}, \beta_{1}$, and $\alpha_{2}, \beta_{2}$ respectively, then the solution for initial values $L \alpha_{1}+M \alpha_{2}, L \beta_{1}+M \beta_{2}$ could be found approximately by

$$
\begin{aligned}
& N=L N_{1}+M N_{2} \\
& N^{\prime}=L N_{1}^{\prime}+M N_{2}^{\prime} \\
& G^{\prime \prime}=L G_{1}^{\prime \prime}+M G_{2}^{\prime \prime}
\end{aligned}
$$

The first few entries for each function were then corrected by accurate calculation from the series expansions ; successive approximations to the true values for all $\zeta$ could then be obtained from these by integrating up the approximate values using the Simpson two-strip rule. The successive approximation process converged very rapidly if $\alpha_{1} \beta_{1}$ and $\alpha_{2} \beta_{2}$ were fairly near points.

Having found $a, b$ and computed the solution based on them, the parameters of the asymptotic series were found simply by substitution. The asymptotic series hold when $\left(K-N^{\prime}\right)^{2}, G^{2}$ are small enough to be ignored ; i.e., for three decimal accuracy, for $\zeta$ such that $K-N^{\prime}, G<0.023$. All values of $N^{\prime}, G$ for $\zeta>\zeta_{0}$, where $\zeta_{0}$ defines this limiting condition, provide estimates of the parameters $A, B, C$; using the means of these estimates, the functions $N^{\prime}, G$ were computed from the asymptotic series, and their values compared with the integrated solution. In no case did the difference exceed 0.001 , so that this method appears to be satisfactory.
11. The case $\mu=0$ (i.e., $K=1, \Omega=0$ ), a stationary disc with forced flow, had already been dealt with by Hömann, who had solved equation (6.2) and obtained the result $a=0.6586$ there is of course no angular velocity. However his method is not satisfactory; he produces a series expansion up to terms in $\zeta^{20}$ and then equates this to an asymptotic solution similar to (8.7) at an apparently arbitrary point $\zeta=1 \cdot 8$. The equations thus produced determine the five unknown parameters. The solution thereby obtained has a discontinuity at $\zeta=1 \cdot 8$, immediately obvious on differencing, and indeed to be expected from the method. In fact this value of $a$ leads to a terminal value for $N^{\prime}$ of 1.015 instead of the required 1 . However the correct
value of $a$ could be determined in this case by interpolation for a single variable only (by use of equation (6.2) which now contained only functions of $N$ and its differential coefficients) ; it proved to be $a=0.656$. Using this solution, equation (6.3) was then solved; for given $N, N^{\prime}$ (6.3) is a linear differential equation for $G$, and so determination of $b$ was by linear interpolation. This value of $b(1.075)$ is meaningless in itself, since $G \equiv 0$ for this case, but it is a limiting value for the other solutions. Coefficients for the asymptotic series were obtained as previously described.
$\mu=\infty$ (i.e. $K=0, \Omega=1$ ) had already been treated by Cochran; he obtained the values $a=0.255, b=-0.616, A_{0}=0.934, B_{0}=1 \cdot 208, C=0.886$. The asymptotic series here takes the form (8.8) and more terms of it can be derived, as higher powers of the exponential $\mathrm{e}^{-c \xi}$.
$\mu=\frac{1}{2}, \mu=1, \mu=2$ were treated by the process described above, using trial values suggested by the solutions already obtained. Tables and graphs of the values of $a, b, A, B, C$ and of the variation of $N, N^{\prime}, G$, and $P$ as functions of $\zeta$, are given at the end of this paper.
[Values of the constants and functions were actually obtained using the equations involving the parameters in a less convenient form ; the above choice was only made later, and the results obtained from the previous process have been converted to those which would have been obtained in this way. The converted tables have been checked by integrating up using the Simpson rule].
12. Frictional Moment Coefficient.-The constant $b$ is of especial interest since it is proportional to the viscous torque exerted by the liquid on the disc. The shearing stress opposing motion is

$$
\rho v\left(\frac{\partial v}{\partial z}\right)_{z=0}=\rho \nu \omega v\left(\frac{\eta}{v}\right)^{1 / 2} b
$$

so that the total retarding moment on one side of a disc of radius $s$, neglecting edge effects, is

$$
\begin{equation*}
M=\int_{0}^{s} \rho v \omega\left(\frac{\eta}{\nu}\right)^{1 / 2} b .2 \pi r^{3} d r=\frac{1}{2} \pi \rho(\nu \eta)^{1 / 2} \omega b s^{4} . \quad . . \quad . \quad . \tag{1.1}
\end{equation*}
$$

Writing $S=\pi s^{2}$, the disc area, and $R=s^{2} \omega^{2} / \nu \eta$, the Reynolds number, (12.1) becomes

$$
\begin{equation*}
M=\frac{1}{2} \rho S^{3} \omega^{2} S R^{-1 / 2} b \quad . . \quad . . \quad . . \quad . . \quad . . \quad \text {.. .. } \tag{12.2}
\end{equation*}
$$

so that we can take $b$ as the torque coefficient

$$
\begin{equation*}
b=\frac{R^{1 / 2}}{\rho S s^{3} \omega^{2}} M . \quad \text {. .. .. .. .. .. .. .. } \tag{12.3}
\end{equation*}
$$

Von Kármán and Cochran use a similar form for their special case $K=0$. The variation of this coefficient with $\mu$ is shown in Table 6 and in graph 5.
13. Inflow Velocity.-The constant $C$ has also direct physical significance. The disc has two distinct effects on the velocity of the fluid along the axis; it acts as a centrifugal fan by its rotation and so sucks the fluid towards it, but it also acts as an infinite plane barrier opposing by viscosity the radial flow due to an externally imposed fluid motion. $C$ is a measure of the resultant of these two effects on the fluid's velocity far from the disc; when $C$ is positive, this resultant effect is a suction, indicating that the 'fan' effect is predominant. When $C$ is negative the barrier effect dominates. As would be expected, when the forced flow is of small importance relative to the rotation (i.e., for $K$ zero or small) $C$ is positive ; it decreases to zero as $K$ increases, vanishing for $K=1 / 4$ approximately, and then decreases further to its final negative value when there is no rotation at all (i.e. $K=1, \Omega=0$ ). This variation is shown in Table 6 and graph 6.
14. Boundary Layer Analogue: Displacement and Momientum Thickness.-The solution found is a solution of the exact Navier-Stokes equations with no boundary layer approximations. Cochran has pointed out that, in the case with which he deals, the solution does in fact bear
out the usual boundary layer assumptions, in that the effect of the motion of the disc on the velocity of the fluid is perceptible only in a layer of thickness $\mathrm{O}\left\{(\nu / \omega)^{1 / 2}\right\}$ and the pressure change is $\mathrm{O}(\rho \nu \omega)$ through that thickness. The general case also fulfils these conditions. It is convenient for this purpose to consider a stationary disc with a fluid given a rotation as well as a forced motion against it ; it follows immediately that the solution for this case is given by

$$
\bar{u}=u, \bar{v}=\gamma \omega-v, \bar{v}=w, \frac{1}{\rho}\left(\bar{p}_{0}-\bar{p}\right)=\frac{1}{\rho}\left(p_{0}-p\right)+\frac{1}{2} \omega^{2} r^{2}
$$

where $u, v, w, p$, are the velocity and pressure components of the solution already given. Then the ' main stream ' velocity tangential to the disc is the resultant of components $k r, \omega r$ and so is $\eta r$; and the fluid velocity becomes indistinguishably different from this for some finite $\zeta$, that is, since $z=(\nu / \eta)^{1 / 2} \zeta$, in a distance $\mathrm{O}\left\{(\nu / \eta)^{1 / 2}\right\}$. The pressure change is $\mathrm{O}(\rho \nu \eta)$, so that the assumptions of boundary layer theory are again satisfied.

We can calculate also from this solution analogues to the boundary layer quantities, displacement thickness ( $\delta$ ) and momentum thickness $(\theta)$. They are defined by the relations

$$
\begin{aligned}
U \delta & =\int(U-u) d z \\
U^{2} \theta & =\int u(U-u) d z
\end{aligned}
$$

where $U$ is the main stream velocity, $u$ the velocity in the boundary layer. In our case these relations become

$$
\begin{align*}
& \eta r \delta=\int\left[\eta \gamma-\left\{u^{2}+(\omega \gamma-v)^{2}\right\}^{1 / 2}\right] \cdot d z \quad . \quad . \quad . \quad . \quad .  \tag{14.1}\\
& \eta^{2} \gamma^{2} \theta=\int\left\{u^{2}+(\omega \gamma-v)^{2}\right\}^{1 / 2}\left[\eta \gamma-\left\{u^{2}+(\omega \gamma-v)^{2}\right\}^{1 / 2}\right] d z \quad . . \quad .  \tag{14.2}\\
& u=\eta r N^{\prime}(\zeta), v=\omega r G(\zeta), \zeta=\left(\frac{\eta}{v}\right)^{1 / 2} z \\
& \delta=\left(\frac{\nu}{\eta}\right)^{1 / 2} \int\left[1-\left\{N^{\prime 2}+\Omega^{2}(1-G)^{2}\right\}^{1 / 2}\right] d \xi \quad . . \quad . \quad .  \tag{14.3}\\
& \theta=\left(\frac{\nu}{\eta}\right)^{1 / 2} \int\left\{N^{\prime 2}+\Omega^{2}(1-G)^{2}\right\}^{1 / 2}\left[1-\left\{N^{\prime 2}+\Omega^{2}(1-G)^{2}\right\}^{1 / 2}\right] d \zeta \tag{14.4}
\end{align*}
$$

or since
$(\eta / \nu)^{1 / 2} \delta,(\eta / v)^{1 / 2} \theta$, and the ratio $\delta / \theta=H$ have been calculated for the various $\mu$, and are shown in Table 7 and graph 7. $H$ is usually assumed constant and taken as 2.4 in turbulence calculations; it may be seen that the value here varies from $2 \cdot 294$ to $2 \cdot 153$ as $\mu$ increases from zero to infinity.
15. My thanks are due to Prof. W. G. Bickley, who suggested this problem to me and has helped at every stage of the work.

## REFERENCES

| No. |  |  |  |  |  |  |  | Title, etc. |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| 1 | .. | . | .. | .. | . | .. | .. | .. | Zeitschr. f. angero. Math. u. Mech. 1 (1921), 244-247. |
| 2 | .. | . | .. | .. | .. | .. | .. | .. | Proc. Camb. Phil. Soc. 30 (1934), 365-375. |
| 3 | .. | .. | .. | .. | .. | .. | .. | .. | Zeitschr. f. anger. Math. u. Mech. 16 (1936), 153-164. |

TABLE 1
TABLE 2

| Values of $G, N, N^{\prime}, P$ |
| :---: |
|  |
|  |
| $\zeta$ |
| 0 |

When $\mu=\infty$

$$
K=0
$$

$$
\Omega=1
$$

Values of $G, N, N^{\prime}, P$
er
$\Omega=0.894$

| $\zeta$ | $G$ | $N$ | $N^{\prime}$ | $P$ |
| :---: | :---: | :---: | :---: | :---: |
| 0 | $1 \cdot 000$ | 0 | 0 | 0 |
| $0 \cdot 1$ | 920 | 0.003 | $0 \cdot 064$ | $0 \cdot 130$ |
| $0 \cdot 2$ | 842 | 012 | 119 | 238 |
| $0 \cdot 3$ | 765 | 027 | 166 | 333 |
| $0 \cdot 4$ | 692 | 046 | 207 | 418 |
| 0.5 | 622 | 068 | 242 | 493 |
| $0 \cdot 6$ | 556 | 094 | 272 | 562 |
| 0.7 | 494 | 122 | 298 | 626 |
| 0.8 | 437 | 153 | 320 | 687 |
| 0.9 | 383 | 186 | 339 | 747 |
| 1.0 | 335 | 221 | 355 | 808 |
| $1 \cdot 1$ | 291 | 257 | 370 | 873 |
| 1.2 | 251 | 295 | 382 | 938 |
| $1 \cdot 3$ | 216 | 334 | 392 | 1.007 |
| $1 \cdot 4$ | 184 | 374 | 401 | 1.082 |
| 1.5 | 156 | 414 | 408 | 1.159 |
| $1 \cdot 6$ | 131 | 455 | 415 | 1.244 |
| 1.7 | 109 | 498 | 420 | 1.336 |
| 1.8 | 091 | 540 | 425 | 1.433 |
| 1.9 | 076 | 583 | 429 | 1.538 |
| $2 \cdot 0$ | 062 | 625 | 433 | 1.647 |
| $2 \cdot 1$ | 051 | 668 | 436 | 1.764 |
| $2 \cdot 2$ | 041 | 711 | 438 | 1.887 |
| $2 \cdot 3$ | 033 | 755 | 440 | $2 \cdot 020$ |
| $2 \cdot 4$ | 026 | 800 | 441 | $2 \cdot 162$ |
| $2 \cdot 5$ | 021 | 844 | 443 | 2.311 |
| $2 \cdot 6$ | 017 | 888 | 444 | $2 \cdot 465$ |
| $2 \cdot 7$ | 013 | 932 | 445 | $2 \cdot 627$ |
| 2.8 | 010 | 977 | 445 | 2.799 |
| $2 \cdot 9$ | $008$ | 1.022 | 446 | 2.981 |
| $3 \cdot 0$ | 006 | 1.066 | 447 | $3 \cdot 167$ |

TABLE 3


TABLE 4
Values of $G, N, N^{\prime}, P$
When $\mu=\frac{1}{2}$
$K=0.894$
$\Omega=0.447$

| $\zeta$ | $G$ | $N$ | $N^{\prime}$ | $P$ |
| :--- | :--- | :--- | :--- | :--- |
| 0 | 1 | 0 | 0 | 0 |
| $0 \cdot 1$ | $0 \cdot 898$ | $0 \cdot 006$ | $0 \cdot 112$ | $0 \cdot 224$ |
| $0 \cdot 2$ | 798 | 022 | 214 | 429 |
| $0 \cdot 3$ | 702 | 048 | 307 | 619 |
| $0 \cdot 4$ | 611 | 083 | 391 | 796 |
| $0 \cdot 5$ | 526 | 126 | 466 | 964 |
| $0 \cdot 6$ | 448 | 176 | 533 | $1 \cdot 128$ |
| $0 \cdot 7$ | 377 | 232 | 592 | $1 \cdot 292$ |
| $0 \cdot 8$ | 314 | 294 | 643 | $1 \cdot 459$ |
| $0 \cdot 9$ | 251 | 360 | 688 | $1 \cdot 636$ |
| $1 \cdot 0$ | 210 | 433 | 727 | $1 \cdot 829$ |
| $1 \cdot 1$ | 169 | 507 | 760 | $2 \cdot 034$ |
| $1 \cdot 2$ | 134 | 584 | 788 | $2 \cdot 258$ |
| $1 \cdot 3$ | 105 | 664 | 812 | $2 \cdot 506$ |
| $1 \cdot 4$ | 082 | 746 | 830 | $2 \cdot 773$ |
| $1 \cdot 5$ | 063 | 830 | 846 | $3 \cdot 070$ |
| $1 \cdot 6$ | 047 | 916 | 858 | $3 \cdot 394$ |
| $1 \cdot 7$ | 035 | $1 \cdot 002$ | 868 | $3 \cdot 744$ |
| $1 \cdot 8$ | 026 | $1 \cdot 089$ | 875 | $4 \cdot 122$ |
| $1 \cdot 9$ | 018 | $1 \cdot 176$ | 880 | $4 \cdot 526$ |
| $2 \cdot 0$ | 013 | $1 \cdot 265$ | 885 | $4 \cdot 970$ |
| $2 \cdot 1$ | 010 | $1 \cdot 354$ | 888 | $5 \cdot 443$ |
| $2 \cdot 2$ | 007 | $1 \cdot 443$ | 890 | $5 \cdot 944$ |
| $2 \cdot 3$ | 005 | $1 \cdot 532$ | 892 | $6 \cdot 478$ |
| $2 \cdot 4$ | 003 | $1 \cdot 621$ | 893 | $7 \cdot 041$ |
| $2 \cdot 5$ | 002 | $1 \cdot 710$ | 894 | $7 \cdot 636$ |
| $2 \cdot 6$ | 002 | $1 \cdot 800$ | 894 | $8 \cdot 268$ |
|  |  |  |  |  |

TABLE 5
Values of $N, N^{\prime}, P$

$$
\text { When } \begin{aligned}
\mu & =0 \\
K & =1 \\
\Omega & =0
\end{aligned}
$$

| $\zeta$ | $N$ | $N^{\prime}$ | $P$ |
| :---: | :---: | :---: | :---: |
| 0 | 0 | 0 | 0 |
| $0 \cdot 1$ | 0.006 | $0 \cdot 126$ | $0 \cdot 253$ |
| 0.2 | 025 | 242 | 486 |
| 0.3 | 055 | 349 | 703 |
| $0 \cdot 4$ | 094 | 445 | 908 |
| 0.5 | 143 | 532 | 1-104 |
| $0 \cdot 6$ | 200 | 609 | 1-298 |
| 0.7 | 265 | 677 | $1 \cdot 494$ |
| $0 \cdot 8$ | 335 | 736 | 1.697 |
| $0 \cdot 9$ | 412 | 787 | $1 \cdot 912$ |
| $1 \cdot 0$ | 492 | 830 | $2 \cdot 145$ |
| $1 \cdot 1$ | 577 | 865 | $2 \cdot 397$ |
| $1 \cdot 2$ | 665 | 896 | $2 \cdot 677$ |
| $1 \cdot 3$ | 756 | 920 | $2 \cdot 985$ |
| $1 \cdot 4$ | 849 | 940 | $3 \cdot 322$ |
| 1.5 | 944 | 955 | $3 \cdot 693$ |
| 1.6 | 1.040 | 967 | 4.099 |
| 1.7 | $1 \cdot 138$ | 975 | 4.538 |
| 1.8 | $1 \cdot 235$ | 983 | 5.019 |
| $1 \cdot 9$ | 1.334 | 988 | 5.536 |
| $2 \cdot 0$ | 1.433 | 992 | 6.092 |
| $2 \cdot 1$ | 1.533 | 995 | 6.686 |
| $2 \cdot 2$ | 1.632 | 997 | 7.320 |
| $2 \cdot 3$ | 1.732 | 998 | 7.994 |
| $2 \cdot 4$ | 1.832 | 999 | 8.706 |
| 2.5 | 1.932 | 999 | 9-460 |

TABLE 6
Values of $a, b, A, B, C$
For various $\mu, K$

| $\mu$ | $\infty$ | 2 | 1 | $\frac{1}{2}$ | 0 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $a$ | 0.255 | 0.343 | $0 \cdot 468$ | 0.583 | $0 \cdot 656$ |
| $b$ | -0.616 | $-0.800$ | -0.933 | $-1.025$ | $-1.075$ |
| A | $A_{0}=0.934$ | 0.133 | 0.269 | 0.358 | $0 \cdot 408$ |
| $B$ | $B_{0}=1.208$ | 0.567 | 0.485 | $0 \cdot 466$ | $0 \cdot 421$ |
| C | 0.886 | -0.275 | --0.440 | $-0.526$ | -0.569 |
| K | 0 | 0.447 | 0.707 | 0.894 | 1 |

TABLE 7

Values of Displacement and Momentum Thickness Coefficients $(\eta / v)^{1 / 2} \delta,(\eta / v)^{1 / 2} \theta$ and $H=\delta / \theta$ for various $\mu, K$

| $\mu$ | $\infty$ | 2 | 1 | $\frac{1}{2}$ | 0 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $(\eta / \nu)^{1 / 2} \delta$ | 1.143 | 0.787 | 0.661 | 0.596 | 0.569 |
| $(\eta / \nu)^{1 / 2} \theta$ | 0.531 | 0.348 | 0.289 | 0.260 | 0.248 |
| $H=\delta / \theta$ | 2.153 | 2.261 | 2.284 | 2.292 | 2.294 |
| $K$ | 0 | 0.447 | 0.707 | 0.894 | 1 |



Fig. 1. Variation of $N^{\prime}$.


FIg. 2. Variation of $G$.


Fig. 3. Vibration of $N$.


Fig. 4. Variation of $P$.



Fig. 6. Variation of $c$.


Fig. 7. Variation of $\delta, \theta$ and $H$.

## Publications of the <br> Aeronautical $\mathbb{R e s e a r c h}$ Council

ANNUAL TECHNICAL REPORTS OF THE AERONAUTICAL RESEARCH COUNCIL (BOUND VOLUMES)-

1934-35 Vol. I. Aerodynamics. Out of print.
Vol. II. Seaplanes, Structures, Engines, Materials, etc. $405 .(405.8 d$. )
1935-36 Vol. I. Aerodynamics, 3os. (30s. 7 d.)
Vol. II. Structures, Flutter, Engines, Seaplanes, etc. 30s. (30s. 7d.)
1936 Vol. I. Aerodynamics General, Performance, Airscrews, Flutter and Spinning. 4os. (40s. gd.)
Vol. II. Stability and Control, Structurees, Seaplanes, Engines, etc. 505. (50s. 1od.)
1937 Vol. I. Aerodynamics General, Performance, Airscrews, Flutter and Spinning. 4os. (40s. Iod.)
Vol. II. Stability and Control, Structures, Seaplanes, Engines, etc. 60 . (6is.)
1938 Vol. I. Aerodynamics General, Performance, Airscrews. 50s. (5Is.)
Vol. II. Stability and Control, Flutter, Structures, Seaplanes, Wind Tunnels, Materials. 30 s . (30s. 9 d .)
1939 Vol. I. Aerodynamics General, Performance, Airscrews, Engines. 50 . (50s. ix d.)
Vol. II. Stability and Control, Flutter and Vibration, Instruments, Structures, Seaplanes, etc. 63 s. $(64 \mathrm{~s} .2 \mathrm{~d}$.
1940 Aero and Hydrodynamics, Aerofoils, Airscrews, Engines, Flutter, Icing, Stability and Control, Structures, and a miscellaneous section. 50s. (5rs.)
Certain other reports proper to the 1940 volume will subsequently be included in a separate volume.
ANNUAL REPORTS OF THE AERONAUTICAL RESEARCH COUNCIL-2933-34 Is. 6d. (Is. 8d.) 1934-35 Is. 6d. (xs. 8d.)
April I, 1935 to December 3 I, 1936. 4 s . (4s. $4 \%$ )
1937 2s. (2s. 2d.)
1938 Is. $6 d$. ( Is .8 d. )
1939-48 3s. (3s. 2d.)
INDEX TO ALL REPORTS AND MEMORANDA PUBLISHED IN THE ANNUAL TECHNICAL REPORTS, AND SEPARATELY-
April, 1950
R. 8 M M. No. 2600 . 2s. 6 d. ( $2 s_{0} 7 \frac{1}{2} d$. )

INDEXES TO THE TECHNICAL REPORTS OF THE AERONAUTICAL RESEARCH COUNCII-

December 1, 1936 - June 30 , 1939. R. ${ }^{\circ}$ M. No. 1850. 1s. 3 d. (IS. $4 \frac{1}{2}$ d.)
July 1, 1939 - June 30, 1945. R. \& M. No. I950. is. (Is. I $\frac{1}{2} d$. )
July 1, 1945-June 30, $1946 . \quad$ R. \& M. No. 2050 . Is. (rs. $1 \frac{1}{2} d$.)
July 1, i946-- December 31, 1946. R. \& M. No. 2150. Is. 3 d. (is. $4 \frac{1}{2} d$. )
January I, 1947 - June 30 , 1947. R. \& Mi. No. 2250. is. 3 d. (is. 4 $\frac{1}{2} d$. )
Prices in brackets include postage.
Obtainable from
HER MAJESTY'S STATIONERY OFFICE
York House, Kingsway, London, w.c. 2423 Oxford Street, LONDON, w. 1 P.O. Box 569, LONDON, S.E. 1

13 a Castle Street, edingurgu, 2 ist. Andrew's Crescent, cardiff
39 King Street, Manchester, 2 Tower Lane, bristol 1
2 Edmund Street, birmingham, 380 Clichester Street, belfast or through any bookseller.

